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### Key Points:

- We conduct 2D thermomechanical models of subduction with variable sediment thickness and density
- Thick sediments can increase resistance along the subduction interface and decrease slab pull leading to a slower subducting plate
- Sediments can act as a lubricant for long-term subduction, but buoyancy and accretionary wedge development are also important

### Supporting Information:

Supporting Information may be found in the online version of this article.

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## The Role of Sediment Accretion and Buoyancy on Subduction Dynamics and Geometry

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**Abstract** Subducted sediments are thought to lubricate the subduction interface and promote faster plate speeds. However, global observations are not clear-cut on the relationship between the amount of sediments and plate motion. Sediments are also thought to influence slab dip, but variations in subduction geometry depend on multiple factors. Here we use 2D thermomechanical models to explore how sediments can influence subduction dynamics and geometry. We find that thick sediments can lead to slower subduction due to an increase of the megathrust shear stress as the accretionary wedge gets wider, and a decrease in slab pull as buoyant sediments are subducted. Our results also show that larger slab buoyancy and megathrust stress due to thick sediments increase the slab bending radius. This offers a new perspective on the role of sediments, suggesting that sediment buoyancy and wedge geometry also play an important role on large-scale subduction dynamics.

**Plain Language Summary** At subduction zones, an oceanic plate dives into the mantle below another plate. The downgoing plate is usually covered by sediments. These sediments can be carried down to depth along the interface and/or scraped off the top of the downgoing plate and appended to the edge of the upper plate, forming an accretionary wedge. Sediments subducted to depth act as a lubricant, influencing the shear resistance of the interface, and in turn, downgoing plate speed. However, natural data show that slow subduction can be associated with thick sediments. Sediments are also thought to affect the dip angle of the downgoing plate, but subduction geometry is also influenced by other factors. We conducted a numerical modeling study to understand the effect of sediment thickness and density on the downgoing plate speed and dip. We observe that thick sediments on the downgoing plate lead to a slower subduction and a shallower dip, due to the decrease in slab pull and increase of stress along the contact interface associated to a bigger accretionary wedge. Our findings suggest that the effect of sediments might be not limited to the lubrication of the contact interface, but buoyancy and accretionary wedge size also play a role.

## 1. Introduction

The main parameters controlling subduction kinematics and geometry remain poorly understood. Previous work suggested that plate motion depends on the balance between the negative buoyancy of the subducting lithosphere (e.g., Becker & O'Connell, 2001; Conrad & Lithgow-Bertelloni, 2002; Forsyth & Uyeda, 1975) and lithospheric bending, mantle resistance, and shear coupling along the subduction interface (e.g., Buffett & Rowley, 2006; Conrad & Hager, 1999). For moderately strong slabs, the plate interface matters (Conrad & Hager, 1999) and interface rheology has been suggested to control plate speeds (Behr & Becker, 2018). Sediments entering the trench can influence the stress state of the megathrust (e.g., Lamb, 2006). Due to their low friction and/or high fluid pressure (e.g., Bangs et al., 2009; Lamb, 2006; Lamb & Davis, 2003; Saffer & Marone, 2003; Saffer & Tobin, 2011), sediments might lubricate the plate interface. Hence, their presence is expected to speed up plate motion, all else being equal. For example, Lamb and Davis (2003) suggested that a decrease of the interface shear stresses in the frictional regime due to thick trench sediments might result

in the acceleration of convergence rate. Behr and Becker (2018) showed that sediment-lubricated slabs subduct faster than slabs with predominantly mafic material, due to the lower viscosity of the deep interface when abundant sediments subduct.

Considering global observations, the role of sediments remains unclear. Trench sediment thickness seems to be negatively correlated with convergence velocity (e.g., Clift & Vannucchi, 2004) or subduction velocity (Duarte et al., 2015). Slow converging systems are usually associated with sediment accretion (e.g., Clift & Vannucchi, 2004). However, this relationship at least partially occurs because the time for sediment accumulation is longer if convergence is slow (e.g., Clift & Vannucchi, 2004). Furthermore, the other variables that affect subduction plate speeds (e.g., slab strength and length, overriding plate thickness, and asthenospheric viscosity) vary widely among modern subduction zones, making it difficult to isolate the effect of interface rheology. Challenges also lie in understanding how subducted sediments are partitioned along the interface at shallow and deep levels in accretionary versus erosional margins (cf., Clift & Vannucchi, 2004). Several studies, for example, suggest that even in sediment-starved erosional margins, sediments pile up through underplating deeper along the subduction interface (Agard et al., 2009; Calvert et al., 2011; Delph et al., 2021; Litchfield et al., 2007; Menant et al., 2020; Tewksbury-Christle et al., 2021), which could lead to lubrication despite very low sedimentation rates at the trench.

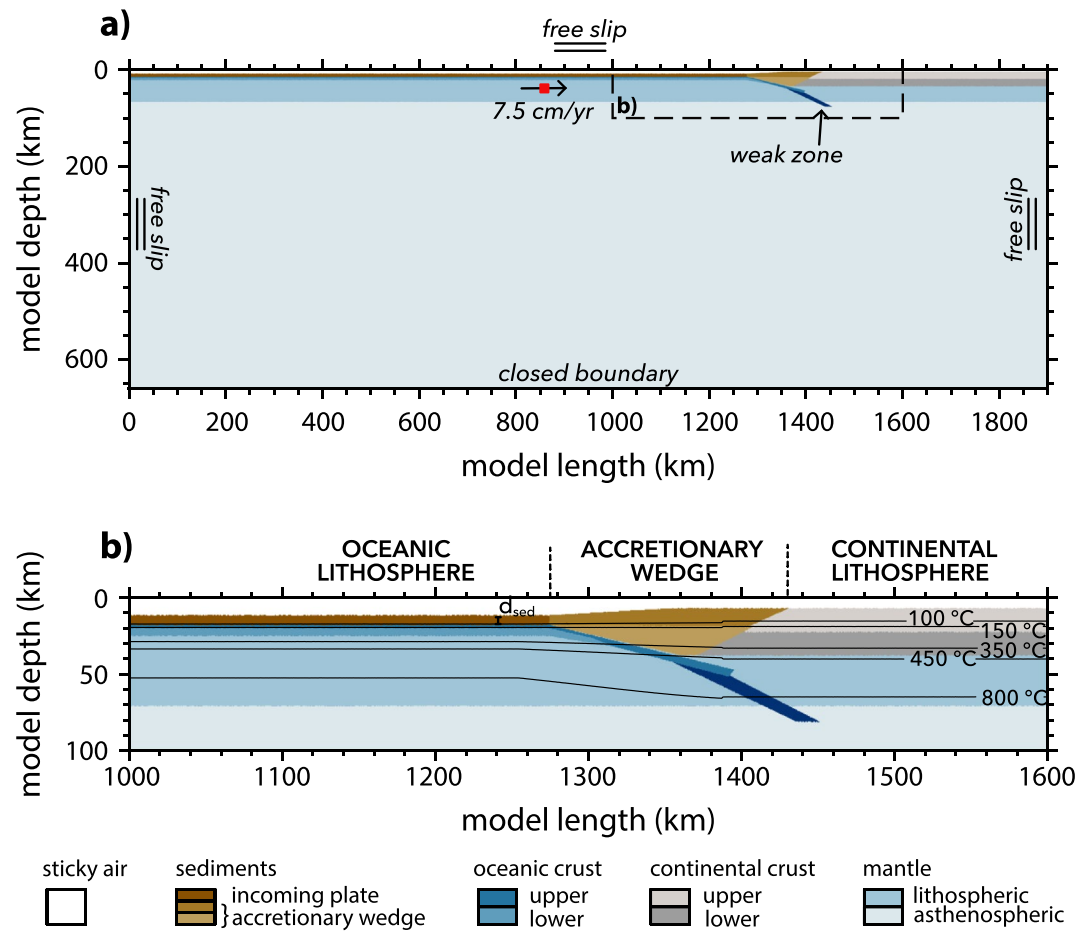
Previous work also focused on the parameters that control the curvature radius of sinking slabs. It has been suggested that slab dip is influenced by a balance between slab buoyancy and hydrodynamic forces related to the corner flow induced in the viscous mantle by the subducting lithosphere (Stevenson & Turner, 1977; Tovish et al., 1978). Trench migration, slab strength, overriding plate thickness and motion with respect to the mantle are also thought to affect the curvature radius of the slab (Bellahsen et al., 2005; Capitanio & Morra, 2012; Capitanio et al., 2009; Funicello et al., 2008; Holt et al., 2015; Lallemand et al., 2005). Numerical models have also suggested that the subducting plate dip can be influenced by sediment thickness at the trench. As the trench sediment thickness increases, the slab unbends due to the seaward growth of the sedimentary wedge (Brizzi et al., 2020).

Here we investigate the role of sediment thickness and buoyancy on the subducting plate velocity and radius of curvature of the slab. We use 2D thermomechanical models in which the slab sinks into the mantle under its negative buoyancy after an initial push. Rather than sediment lubrication, our setup allows us to isolate the effects of sediment buoyancy. We test how the thickness and density of sediments influence slab pull and shear stress at the subduction interface, and we compare these outcomes with slab velocity and curvature radius.

## 2. Numerical Methods, Model Setup, and Model Metrics

We use the 2D Seismo-Thermo-Mechanical version (van Dinther et al., 2013) of the geodynamic code I2ELVIS (Gerya & Yuen, 2007). This solves for the conservation of mass, momentum, and energy using a finite difference scheme on a fully staggered Eulerian grid in combination with Lagrangian markers. Except for the asthenospheric mantle that is Newtonian for simplicity, we employ non-Newtonian visco-elasto-plastic rheologies (Gerya & Yuen, 2007). The effective viscosity is calculated from experimentally constrained dislocation creep flow laws (Table S1 in Supporting Information S1). Incoming plate and accretionary wedge sediments are assumed to be dominantly terrigenous, with a small pelagic component and are modeled using a wet quartzite flow law. The upper oceanic crust is also modeled with a wet quartzite flow law, while the lower oceanic crust is treated as plagioclase. Frictional plastic deformation uses a Drucker-Prager yield criterion (Drucker & Prager, 1952). Differences in the frictional behavior of sediments and oceanic lithosphere are mainly related to the (a) static friction coefficient ( $\mu_s = 0.35$  and  $0.5$  for sediments and oceanic lithosphere, respectively) and (b) pore fluid pressure factor ( $\lambda = 0.95$  and  $0.4$  for sediments and oceanic lithosphere, respectively). Incoming plate and accretionary wedge sediments have dominantly plastic behavior for  $T \sim \leq 350$  °C (i.e.,  $\sim \leq 50$ – $60$  km depth), while material subducted deeper along the plate interface deforms viscously (Figure S1 in Supporting Information S1).

We adapt the model setup of Brizzi et al. (2020), which was originally designed to resemble the South Chile margin (Gorczyk et al., 2007). A 40 Myr old oceanic lithosphere that includes a sedimentary layer of variable thickness  $d_{sed}$  subducts beneath a continental lithosphere (Figure 1). A sedimentary wedge is present at



**Figure 1.** (a) Model setup. Subduction proceeds along a weak zone by imposing a fixed velocity (black arrow) on a small region (red rectangle) of the subducting plate until 300 km of oceanic lithosphere has subducted. The dashed black rectangle marks the high resolution area of the models. (b) Zoom of the high resolution area. Black solid lines show the initial temperature field.  $d_{sed}$  is the thickness of incoming plate sediments (set to 6 km in this model).

the leading edge of the overriding plate, as imaged by seismic profiles and geophysical data (Bangs & Cande, 1997). A 12.5 km thick layer of sticky air mimics the effect of a free surface (e.g., Cramer et al., 2012). Free slip boundary conditions are applied at the top and side boundaries of the model, and we impose a closed boundary condition at the bottom boundary. Subduction initiates along a 15°-dipping weak zone (low plastic strength). We impose a constant velocity of 7.5 cm/yr until 300 km of the slab is subducted. After this kinematically prescribed phase, the pushing velocity is removed and subduction is self-driven. An extended description of the numerical methodology and model setup is given in Supporting Information S1.

For each model, we measure the area of subducted sediments  $d_{ss}$ , slab velocity  $v_{sp}$  during the free sinking phase, radius of curvature  $R_c$ , slab pull  $F_{sp}$  and integrated shear stress along the megathrust  $F_{sl}$ .  $d_{ss}$  is defined as the area of sediments below the continental Moho (Figure S2b in Supporting Information S1).  $v_{sp}$  is defined as the average velocity of the subducting plate during the free sinking phase, i.e., from ~4 Myr until the slab reaches the 660 km discontinuity (Figure S3a in Supporting Information S1).  $R_c$  is estimated by fitting a circle to the subducting plate center line.  $F_{sp}$  (force per length) is defined as  $F_{sp} = \Delta \rho g A$ , where  $\Delta \rho$  is the density contrast between the asthenospheric mantle and the slab,  $g$  is the gravitational acceleration, and  $A$  is the slab area (Figure S4 in Supporting Information S1). We compute  $F_{sp}$  at the beginning of self-consistent subduction (~4 Myr) to ensure that an equal length of slab has subducted in each model. Lastly,  $F_{sl}$  is quantified from the length-integrated second invariant of the deviatoric stress tensor in a 3 km-thick region that extends from the trench to the brittle-ductile transition (~450°C isotherm; Figure S5 in Supporting Information S1). To be able to compare with slab pull estimates,  $d_{ss}$ ,  $R_c$ , and  $F_{sl}$  are also measured at ~4 Myr.

### 3. Results

We investigate how sediments influence subduction by varying their (a) thickness  $d_{sed}$  from 0 to 6 km and (b) density  $\rho_{sed}$  from 2,200 to 2,800 kg/m<sup>3</sup>. Note that we vary the density of both incoming plate and accretionary wedge sediments. In the following, we first present the evolution of the models with no ( $d_{sed} = 0$  km) and a thick ( $d_{sed} = 6$  km) sediment layer on the incoming plate and a reference  $\rho_{sed}$  of 2,800 kg/m<sup>3</sup>. Then, we address the evolution of the model with thick light sediments ( $d_{sed} = 6$  km and  $\rho_{sed} = 2,200$  kg/m<sup>3</sup>). Lastly, we focus on the effect of sediments on slab velocity and curvature radius.

#### 3.1. Model Evolution

##### 3.1.1. No Sediment Layer

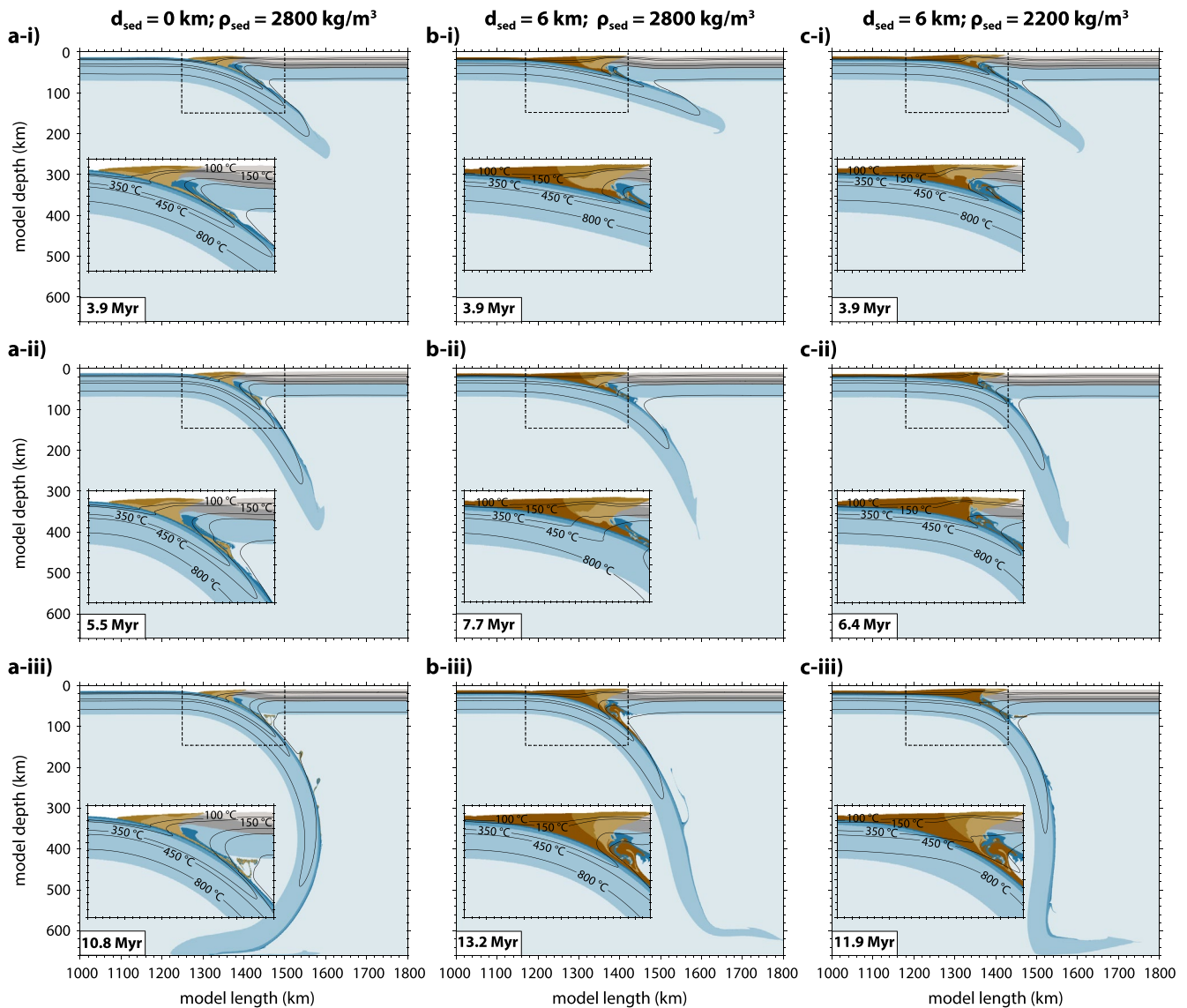
During the initial phase of forced subduction, sediments are eroded from the pre-existing accretionary wedge and transported along the interface up to a maximum depth of ~80 km within a thin subduction channel (Figure 2a-i). When we stop pushing the subducting plate, the slab dip increases (Figure 2a-ii). Sediments are still eroded from the accretionary wedge and transported to a maximum depth of ~100 km along the megathrust (Figure 2a-ii). During this stage, slab velocity increases (Figure S3a in Supporting Information S1) due to both an increase of slab pull and a decrease of the integrated shear resistance at the base of lithospheric mantle. With ongoing subduction, the slab steepens and becomes almost vertical. When it approaches the 660 km discontinuity (i.e., bottom model boundary), the slab tip is slightly overturned. This promotes a backward reclined configuration with progressing subduction (Figure 2a-iii). Sediments subducted below the forearc mantle wedge (depth > 100 km) start detaching and exhuming below the continental lithosphere.

##### 3.1.2. Thick Sediment Layer

During the kinematically prescribed subduction, sediments are partially subducted along the megathrust and partially accreted. Accretion occurs both by off-scraping at the front of the wedge and underplating at the rear. The maximum depth reached by subducted sediments is lower compared to the no sediment model (Figure 2b-i), as underplating promotes the development of an antiformal stack within a thick subduction channel. The dip angle of the slab is lower compared to the no sediment model (Figure 2b-i). As the pushing velocity is removed, sediments keep piling up onto the base of the accretionary wedge, while a small amount is subducted below the continental Moho (Figure 2b-ii). Subduction maintains a shallower dip compared to the no sediment case (Figure 2b-ii). During this phase, slab velocity increases but to a lower rate compared to the no sediment case (Figure S3a in Supporting Information S1). As the slab approaches the 660 km discontinuity, the dip angle increases. This change in the slab dip promotes an increase of the subduction channel width, such that a larger amount of sediments can be dragged to greater depths and underplate onto the base of the accretionary wedge (Figure 2b-iii). During sinking, the slab stretches and eventually drapes over the 660 km discontinuity (Figure 2b-iii).

##### 3.1.3. Thick Light Sediments

During the initial phase of forced subduction, a low sediment density favors more sediment accretion than subduction. Therefore, at the end of the forced subduction, the amount of sediments below the forearc Moho is lower compared to the reference model. This is because the lower density inhibits sediment descent into the subduction channel (Figure 2c-i). At this stage, the slab dip is slightly higher than the respective reference model (Figure 2c-i). As the slab sinks freely into the mantle, the amount of sediments accreted to the wedge increases, while the amount of subducted sediments decreases (Figure 2c-ii). During this stage, the slab dip increases. As observed for the respective reference model, this increase in slab dip induces an increase of the subduction channel width, hence an increase of the amount of subducted sediments. However, with ongoing subduction, these sediments tend to be transported upward to the opening of the channel (Figure 2c-iii). Slab velocity increases as well, but to a higher rate compared to the respective reference model (Figure S3a in Supporting Information S1). As the slab approaches the 660 km discontinuity and drapes over it, significant underplating below the continental lithosphere occurs and a sub-horizontal sedimentary plume develops (Figure 2c-iii).



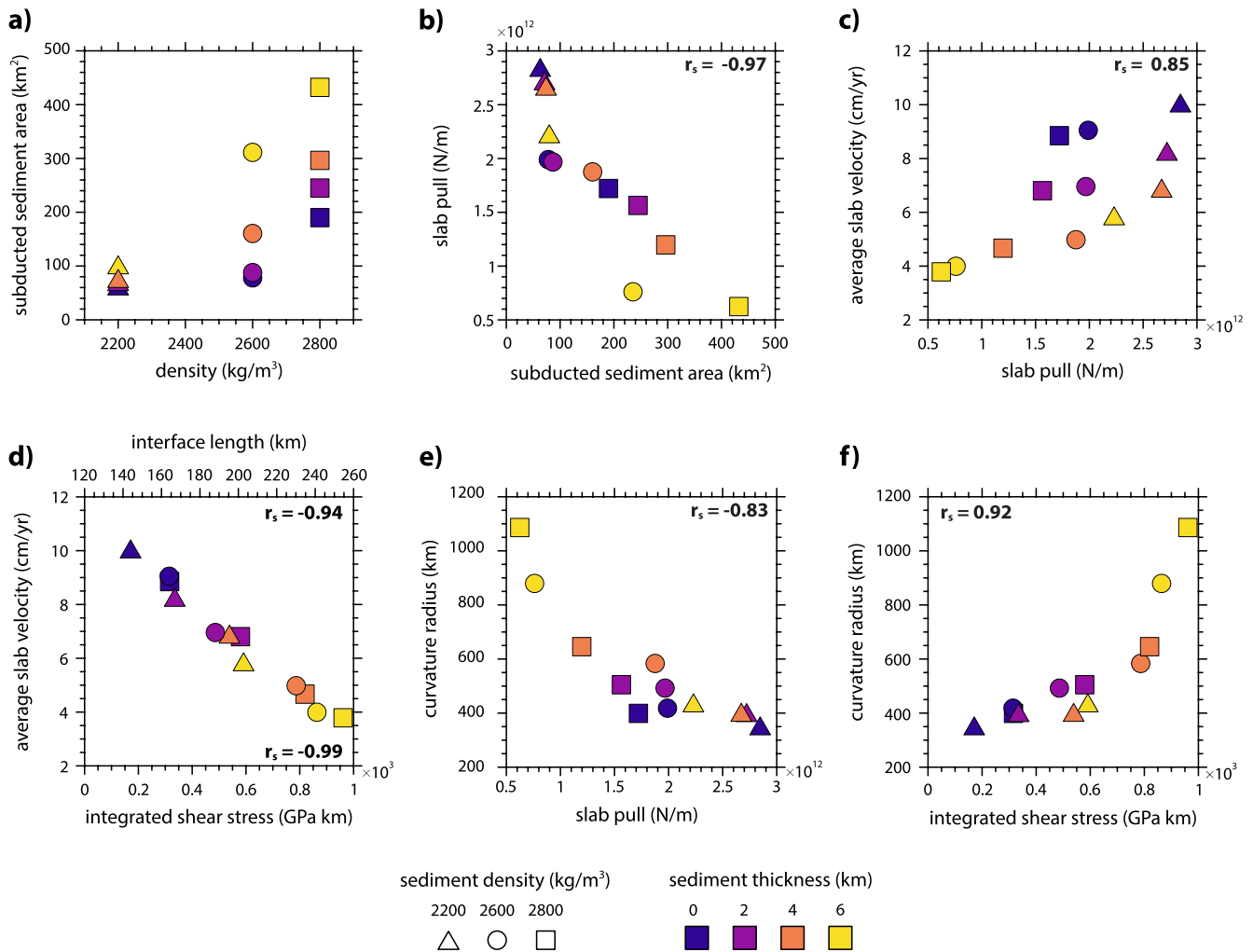
**Figure 2.** Compositional map of the model with (a) thin sediments ( $d_{sed} = 0$  km;  $\rho_{sed} = 2,800$  kg/m<sup>3</sup>), (b) thick sediments ( $d_{sed} = 6$  km;  $\rho_{sed} = 2,800$  kg/m<sup>3</sup>), and (c) thick light sediments ( $d_{sed} = 6$  km;  $\rho_{sed} = 2,200$  kg/m<sup>3</sup>) roughly at the end of the (i) kinematically constrained subduction, (ii) free slab sinking, and (iii) interaction with the 660 km discontinuity. Black lines correspond to 100°C, 150°C, 350°C, 450°C, and 800°C isotherms. Color legend for rock types in Figure 1.

### 3.2. Sediment Control on Slab Velocity

Our results show that the amount of subducted sediments depends on their initial thickness and density (Figure 3a). An increase of sediment thickness results in an increase of subducted sediments. For example, for a sediment density of 2,800 kg/m<sup>3</sup>,  $d_{ss}$  increases by a factor of  $\sim 2$  when  $d_{sed}$  is increased from 0 to 6 km. For a constant sediment thickness, decreasing sediment density results in a decrease of the amount of material subducted, as a relatively higher sediment buoyancy inhibits subduction. For example, if  $d_{sed} = 0$  km,  $d_{ss}$  decreases by a factor of  $\sim 3$ , if  $\rho_{sed}$  decreases from 2,800 kg/m<sup>3</sup> to 2,200 kg/m<sup>3</sup>. This decrease is higher (factor of  $\sim 5.4$ ) if  $d_{sed} = 6$  km.

The amount of subducted sediments influences slab pull  $F_{sp}$  (Figure 3b). As the sediment thickness increases and more sediments are subducted,  $F_{sp}$  decreases by a factor of  $\sim 1.2$  and  $\sim 2.7$  for  $\rho_{sed}$  of 2,800 kg/m<sup>3</sup> and 2,200 kg/m<sup>3</sup>, respectively. As we decrease  $\rho_{sed}$  and the amount of subducted sediments decreases,  $F_{sp}$  increases by a factor of  $\sim 1.6$  and  $\sim 3.6$ , if  $d_{sed}$  is 0 and 6 km, respectively.





**Figure 3.** (a) Relationship between sediment thickness, sediment density and amount of sediments subducted below the forearc Moho; (b) slab pull as a function of the amount of subducted sediments; (c) slab velocity as a function of slab pull; (d) slab velocity as a function of megathrust integrated shear stresses and interface length; (e) radius of curvature the slab as a function of slab pull; (f) curvature radius of the slab as a function of megathrust integrated shear stresses.  $r_s$  is Spearman's rank correlation coefficient. In panel (d) the top and bottom  $r_s$  values refer to the relationship between average slab velocity and interface length, and average slab velocity and integrated shear stress, respectively.  $p$ -values of all relationships is  $< 0.05$ .

Subducting plate velocity  $v_{sp}$  is positively correlated to slab pull (Figure 3c). For  $\rho_{sed} = 2,800 \text{ kg/m}^3$ , the decrease in slab pull that results from the increase in sediment thickness causes a decrease in  $v_{sp}$  from 8.8 to 3.8 cm/yr. On the other hand, as slab pull increases due to a decrease of sediment density,  $v_{sp}$  increases. For example, the increase in slab pull observed when  $d_{sed} = 6 \text{ km}$  and  $\rho_{sed}$  decreases from  $2,800 \text{ kg/m}^3$  to  $2,200 \text{ kg/m}^3$  results in an increase of  $v_{sp}$  from 3.8 to 5.9 cm/yr.

We test how the initial kinematically imposed subduction affects slab velocity by pushing the subducting plate at lower rates. We find that a lower pushing velocity results in a slower slab only in the case of thick sediments (Figure S3b in Supporting Information S1) due to an increase in interface viscosity, as we remove the push and the strain rate decreases. Nonetheless, thick trench sediments result in a slower subducting plate.

Our results also show that increasing the sediment thickness produces an increase of the integrated megathrust shear stress  $F_{sl}$  by a factor of  $\sim 2$  and  $\sim 2.2$ , if  $\rho_{sed}$  is  $2,200 \text{ kg/m}^3$  and  $2,800 \text{ kg/m}^3$ , respectively (Figure 3d). This increase is mainly due to the development of a wider accretionary wedge that increases the interface downdip length (Figure 3d). As  $F_{sl}$  increases,  $v_{sp}$  decreases (Figure 3d). As opposed to the effect of  $d_{sed}$ , a decrease in density promotes a decrease of  $F_{sl}$  by a factor of  $\sim 1.4$ – $1.5$ , hence an increase of  $v_{sp}$  (Figure 3d).

### 3.3. Sediment Control on Slab Curvature Radius

Our results show that sediment thickness and density also influence the curvature radius of the slab  $R_c$ . We find that there is a positive relationship between slab pull and  $R_c$ . As  $F_{sp}$  decreases with increasing  $d_{sed}$  (Figure 3b) and the slab gets more buoyant, subduction attains a flatter geometry and the curvature radius increases by a factor of  $\sim 1.2$  and  $\sim 2.7$ , if  $\rho_{sed}$  is  $2,800 \text{ kg/m}^3$  to  $2,200 \text{ kg/m}^3$ , respectively (Figure 3e). Conversely, when  $F_{sp}$  is higher due to lighter sediments,  $R_c$  is  $\sim 1.1$ – $2.7 \times$  lower (Figure 3e) and we observe a steeper dip angle.

We also observe a positive correlation between the slab curvature radius and the megathrust shear stress. As  $F_{sl}$  increases due to an increase in  $d_{sed}$ ,  $R_c$  increases (Figure 3f). As  $F_{sl}$  decreases due to a lower  $\rho_{sed}$ ,  $R_c$  decreases (Figure 3f).

## 4. Discussion

### 4.1. Sediments and Slab Velocity

Sediment subduction is thought to impact plate motion at convergent margins (e.g., Behr & Becker, 2018; Lamb & Davis, 2003). This hypothesis relies upon the notion that subducted sediments influence the shear strength of the megathrust (e.g., Lamb, 2006; Lamb & Davis, 2003). Given their weakening and/or lubricating effect on the plate interface (e.g., Saffer & Marone, 2003; Saffer & Tobin, 2011), subducted sediments are thought to favor higher plate speed (e.g., Behr & Becker, 2018; Lamb & Davis, 2003).

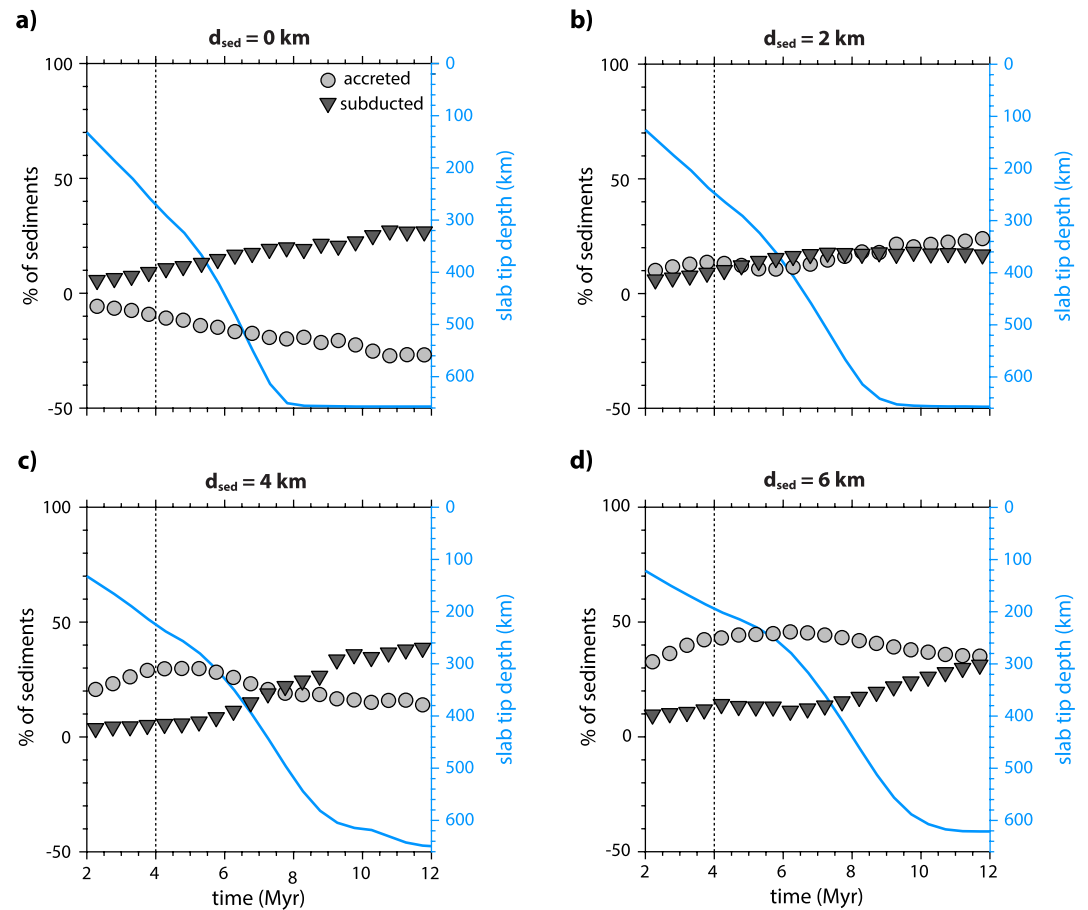
Our results show that slab velocity is indeed affected by interface stress and that a negative correlation between slab velocity and integrated megathrust shear stress exists (Figure 3d), as expected from force balance (Conrad & Hager, 1999). However, we show that as the incoming sediment thickness and density increase, the integrated shear stress along the megathrust increases as well (Figure 3d). Given that shear stress averaged over the megathrust does not vary significantly as a function of sediment thickness and density (Figure S6a in Supporting Information S1), this increase is mainly related to an increase of the interface length (Figure 3d, Figure S6b in Supporting Information S1) due to the presence of a wider accretionary wedge that thickens the upper plate. The larger interface length promotes an increase of the total resistance to subduction, which eventually slows down the slab (Figure 3d).

Subducted sediments decrease plate speed also by decreasing slab pull due to their positive buoyancy. We find that increasing the incoming plate sediment thickness favors the formation of a thick subduction channel, and a large amount of sediments can be subducted (Figure 3a) resulting in a reduction of slab pull (Figure 3b) and, in turn, lower subduction velocity (Figure 3c). Keum and So (2021) showed that sediment buoyancy affects trench motion, with thick trench sediments resulting in a slower trench retreat. This relationship between amount of subducted sediments, slab pull and velocity is also supported by the outcomes of models with different sediment density. Low sediment density makes sediment subduction more difficult, so that slab pull is higher if sediments are relatively light (Figure 3b). This causes higher slab velocities for such lower sediment densities (Figure 3c).

Based on a global compilation, Clift and Vannucchi (2004) suggested that convergence rate is correlated with the thickness of trench sediments because for a constant sediment supply, slower convergence increases the time period over which sediments can accumulate in the trench. Our results suggest, however, that convergence rate is itself linked to sediment supply. Furthermore, our models demonstrate a case in which, for constant interface rheology, sediment supply affects convergence velocity by modulating buoyancy of the subducting lithosphere and the interface length through the construction of an accretionary wedge.

### 4.2. Sediments and Slab Curvature Radius

Our results show that larger integrated megathrust shear stresses result in a larger slab curvature radius (Figure 3f) due to the development of a wide accretionary wedge that increases the interface downdip width. This is in agreement with previous studies suggesting that accretion of sediments can load and unbend the slab, reducing the angle of subduction (Brizzi et al., 2020; Cross & Pilger, 1982; Jacob et al., 1977; Karig & Sharman, 1975; Seely et al., 1974). Comparison with global observations at subduction zones in-



**Figure 4.** Percentage of accreted and subducted sediments, and slab tip depth as a function of time for (a)  $d_{sed} = 0$  km and, (b)  $d_{sed} = 2$  km, (c)  $d_{sed} = 4$  km, and (d)  $d_{sed} = 6$  km. Sediment density  $\rho_{sed}$  is  $2,800 \text{ kg/m}^3$ . Note that the percentage of accreted sediments in panel (a) is negative due to the decrease in size of the proto-wedge as sediments are eroded. The dashed black line marks the timing of push removal. Details on how the percentage of accreted and subducted sediments are estimated are given in Supporting Information S1.

deed suggests that sediment-rich margins are often associated with more shallowly dipping slabs (e.g., Brizzi et al., 2020; Clift & Vannucchi, 2004). Similarly, thick overriding plates have been shown to increase the curvature radius of the slab (Capitanio et al., 2011; Holt et al., 2015).

We also find that there is a negative relationship between slab pull and slab curvature radius (Figure 3e). With increasing subducted sediments, slab pull decreases (Figure 3b) and subduction attains a shallower dipping geometry. Slab dip is expected to be influenced by slab pull (e.g., Molnar & Atwater, 1978; Uyeda & Kanamori, 1979; Vlaar & Wortel, 1976). However, analog experiments show that a larger slab pull promotes slab rollback and shallowing (Funciello et al., 2008; Martinod et al., 2005). Furthermore, a correlation between subducting plate age and slab dip (Cruciani et al., 2005; Lallemand et al., 2005) or slab pull force (Lallemand et al., 2005) is not found in compilations of natural subduction zone parameters. Our findings confirm that the overriding plate structure can influence subduction geometry, but also suggest that slab pull force might factor in.

#### 4.3. Sediment Accretion versus Subduction

It is widely recognized that subduction zones can either be accretionary or erosive (e.g., von Huene & Scholl, 1991), but the mechanisms by which sediments are subducted/eroded or accreted are still debated. Our results confirm previous suggestions that the amount of sediments influences whether accretion or erosion occurs (Figure 4) (e.g., Clift & Vannucchi, 2004; Cloos & Shreve, 1988; von Huene & Scholl, 1991). In



our models, the lack of sediments results in the erosion and subsequent subduction of the wedge sediments (Figure 4a). As the sediment thickness increases, sediments are mostly accreted to the front of the wedge (Figures 4b–4d), but sediment underplating also occurs. However, this is only observed when the sediment thickness is larger than 4 km, suggesting that the amount of sediments can influence whether this process takes place. Based on geophysical observations, underplating is indeed suspected at the Hikurangi (Bassett et al., 2010), Nankai (Kimura et al., 2010), Cascadia (Calvert et al., 2011) and Alaska (Moore et al., 1991) margins, which all have a relatively high trench sediment thickness (Clift & Vannucchi, 2004). Our models also show that sediment subduction takes place along with accretion. Subducted sediments deeper than the megathrust seismogenic zone have not been imaged directly, but their isotopic signatures are found in arc magmas (e.g., Plank & Langmuir, 1998). In our models, the maximum depth to which sediments are subducted seems to depend on their initial thickness. If this is larger than 4 km, sediments can reach ~120 km depth before detaching from the subducting plate and then buoyantly rise to underplate, forming a horizontal plume that stagnates below the continental lithospheric mantle (Figure 2b). Rheological properties are also expected to influence the behavior of subducted sediments. Currie et al. (2007) showed that for sediments with wet quartzite rheology, sediment density exerts the primary control on whether sediment subduction can occur. As sediment viscosity increases, entrainment by the subducting plate tends to dominate and sediments are more easily subducted to depth (Currie et al., 2007).

Convergent margins with high sediment supply are also commonly considered loci of sediment accretion (e.g., Clift & Vannucchi, 2004; Cloos & Shreve, 1988), but transitions to an erosional regime have been documented in Costa Rica, northern Apennines and southern Alaska (Amato & Pavlis, 2010; Vannucchi et al., 2004, 2008). The triggers for switching from one tectonic regime to another remain poorly known. Our models show that the increase in slab dip during the free subduction phase allows for the widening of the subduction channel, such that the amount of subducted sediments increases through time (Figures 4c and 4d). Due to the increase of the subduction channel capacity, the accretionary wedge dynamically readjusts after attaining steady state conditions (e.g., Willett & Brandon, 2002), such that the amount of accreted sediments decreases to accommodate the increase in sediment subduction (Figures 4c and 4d). Thus, the partition of the incoming plate sediments in accreted or subducted is a time-dependent feature, which seems to be strongly influenced by the slab dip (Cloos & Shreve, 1988). Hence, the common view of accretionary or erosive convergent margins seems to be overly simplified (e.g., Simpson, 2010), as sediment subduction and accretion are interlinked processes.

#### 4.4. Modeling Limitations

Our initial geometry includes a proto-accretionary wedge that has been shown to influence both slab velocity and radius of curvature by influencing the integrated megathrust shear stress. Results of additional tests show that the slab subducts slightly faster and steeper if the initial size of the accretionary wedge is smaller. While this reemphasizes the importance of the initial conditions on model behavior, it also confirms more generally our finding that interface stresses and overriding plate structure have significant control on subduction dynamics. All models have the same effective interface rheology. Even for our sediment-starved case, a small amount of sediments is present along the interface due to the erosion of the initial accretionary wedge. This means that our models do not allow us to investigate trade-offs between the degree of sediment lubrication and sediment buoyancy, but instead they effectively isolate the buoyancy effects. The constant sediment flux to the trench in our models does not fully capture variations in sediment thickness in nature. Our slab pull estimates are low compared to the typical values of  $10^{13}$  N/m (e.g., Turcotte & Schubert, 2002), as we derive them at the initial stage of subduction. This suggests that for a young (40 Myr old) plate, sediment buoyancy has a pronounced effect, but we caution that this effect might be lower for older, thicker lithosphere. Eclogitization of the mafic components is expected to contribute to slab pull and influence the force balance, as well the integrated shear stress (Behr & Becker, 2018). Additional aspects that we neglect are fluid transport and compaction effects, as well metasomatic alteration of subducting sediments (Saffer & Tobin, 2011). Erosion and sedimentation are not included in our models, but we might expect these processes to influence both slab velocity and curvature radius by affecting the sediment supply to the trench. Our simulations are 2D, and so we neglect along-strike variations of subducted sediments, which are shown to be important for along-strike variations of trench velocity and curvature (Keum & So, 2021). Despite

such simplifications, our numerical models allow us to identify important effects of sediment thickness and buoyancy on slab dynamics and to better understand long-term behavior of convergent margins.

## 5. Conclusions

Sediment subduction can affect the interface geometry and effective slab pull, hence slab morphology and subducting plate speed. Thick sediments promote thickening of the overriding plate through the development of a wide accretionary wedge that increases the downdip length of the plate interface, hence resistance to subduction. Thick sediments can also slow down the subducting plate by partly offsetting the negative buoyancy of the slab. The larger integrated interface shear stress and slab buoyancy due to thick sediments promote a larger curvature radius of the slab. Accretionary margins can experience periods of erosion due to changes in the slab dip that can result in oscillations of subduction rate and megathrust stress over time. We suggest that the effect of sediments on subduction dynamics is not straightforward. Future studies should address not only the capacity of sediments to lubricate and/or weaken the plate interface, but also how their presence affects wedge and subduction dynamics.

## Data Availability Statement

Model executables, input and output files for the model with  $d_{sed} = 0$  km and  $d_{sed} = 6$  km, and  $\rho_{sed} = 2,800$  kg/m<sup>3</sup> are available on Zenodo (<https://doi.org/10.5281/zenodo.5551171>).

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